ERNST ABBE'S Theory of Image Formation in the Microscope

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Written and published by Otto Lummer and Fritz Reiche under the title Die Lehre von der Bildentstehung im Mikroskop von Ernst Abbe Translated and annotated, with additional material, by Anthony Yen and Martin Burkhardt

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Introduction

Geometrical optics assigns reality to light rays and assumes that where light rays intersect is also where light concentration actually occurs. This arithmetic optics seeks accordingly to evaluate optical systems in such a way that two spaces are imaged onto each other point for point; i.e., outgoing rays from one point in one space (object space) reunite at one point in another space (image space). If an optical system meets this condition, it then transforms the outgoing convex spherical wavefront from the object point to a concave spherical wavefront whose center is the image point. Arithmetic optics does not have to deliver any more than this.

In order to understand the actual light distribution in the center of the concave spherical wavefront, i.e., the image point, image formation must be handled based on wave theory as a diffraction problem. One usually expresses the result of this approach by overlaying on the point of convergence of the homocentric ray bundle (image point in geometrical optics) the diffraction phenomenon that is uniquely determined by the type of blocking to the spherical wavefront in the image space. In reality, the process is reversed: the diffraction phenomenon is the primary image-forming process, and the image point is secondary. In fact, the image where the imaging ray bundle is limited by a circular aperture is at best a diffraction disk with alternating dark and bright diffraction rings of rapidly decreasing intensity. The greater the image angle whose sine is given by the ratio of the radius

Imaging laws of geometrical Opticsⁱⁱ

§1. Construction of a ray refracted by a spherical surface

Let M (Fig. 1) be the center of the refracting sphere of radius r and refractive index n', and the ambient medium have the refractive index n. To find the refracted ray from the incident ray LE, we insert, according to the elegant method of construction of Weyerstraß, two auxiliary circles 1 and 2 with radii

$$r_1 = \frac{n'}{n}r$$

and

$$\mathbf{r}_2 = rac{\mathbf{n}}{\mathbf{n}'}\mathbf{r}$$
 ,

extend ray LE until it intersects auxiliary circle 1 at A, and connect E with point A' where line AM and auxiliary circle 2 intersect. Line EA'L' is the refracted ray associated with LE.

From the similarity of triangles EAM and EA'M, it follows that

$$\angle MEA = \angle EA'M^{iii}$$

that follows, we have the theorem: the object space is imaged point-topoint in the image space. Planes perpendicular to the axis in the object space correspond point-to-point to the planes perpendicular to the axis in the image space.

If one applies the Lagrange relation to each refracting surface in the system successively, one obtains the Lagrange–Helmholtz relation

$$\left. \begin{array}{c} \beta \cdot \gamma = \frac{n}{n'} \\ \text{or} \qquad y'n' \tan u' = yn \tan u \end{array} \right\} , \qquad (5)$$

where β and γ now denote the lateral magnification and angular magnification with respect to the *entire system*, and n and n' are refractive indices of the front (object) and back (image) media.

§5. Imaging equations according to Abbe

In Fig. 6, let there be conjugate pairs of planes L and L' as well as Q and Q', and the associated lateral magnifications be given by

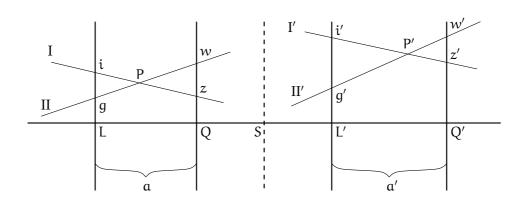


Figure 6

Imaging of self-luminous objects in terms of wave theory

§7. Diffraction problems solved on the basis of Maxwell's theory We have seen that a centered system (microscope objective) images a surface element point-to-point and in similarity, using arbitrarily wide-angled ray bundles, only if the sine condition

$$\frac{\sin u'}{\sin u} = \frac{n}{n'} \cdot \frac{1}{\beta}$$

is fulfilled. If the system is so designed that this condition is satisfied, then all incoming rays to any point of the image remain perpendicular to a spherical surface centered on this point.^{xvi} The lens designer^{xvii} cannot offer anything more than this. We wonder whether and under what conditions this purely geometrical, pointwise concentration of rays is also physically present. Let us for the moment remain on the fiction of geometrical optics, that there were actually luminous points, so only the spherical wave emanating from this point would be a reality. Only with free, absolutely unhindered propagation, as would be the case in an arbitrarily extended, homogeneous medium,

trical waves (Großmann⁹). Finally, the diffraction phenomenon on metallic cylinders of elliptical cross section was treated (B. Sieger¹⁰ and K. Aichi¹¹), if only for material of infinitely large conductivity.

§8. The Kirchhoff principle

In general, the treatment of diffraction phenomena according to the Kirchhoff principle gives a far simpler form, allowing then the calculation of cases of our interest. Applying Green's theorems^{xviii} to a function φ , which satisfies the wave equation^{xix}

$$\frac{\partial^2 \varphi}{\partial t^2} = a^2 \Delta \varphi , \qquad (12)$$

Kirchhoff¹² obtained the value of the function φ at an observation point P (Fig. 11) as a function of time t in terms of values of φ , $\partial \varphi / \partial t$, and $\partial \varphi / \partial \nu$ on the observation point–enclosing surface Σ with inward normal ν ; here one must, for the magnitudes of φ , $\partial \varphi / \partial t$, and $\partial \varphi / \partial \nu$, insert the values that they possess at position d σ at time t' = t - r/a, where r denotes the radius vector P d σ and a the velocity of light in space V. It is^{xx}

$$\varphi_{P}(t) = \frac{1}{4\pi} \int_{\Sigma} d\sigma \left[\varphi \frac{\partial (1/r)}{\partial \nu} - \frac{1}{\alpha r} \frac{\partial \varphi}{\partial t} \cdot \frac{\partial r}{\partial \nu} - \frac{1}{r} \frac{\partial \varphi}{\partial \nu} \right]_{t'=t-\frac{r}{\alpha}} .$$
(13)

Kirchhoff used this theorem to derive an approximation of the light intensity at observation point P (Fig. 12), if waves originating from L are disturbed by some obstacles. We want to carry out the calculation for the special case of an obstacle that is an *opaque* screen with aperture Σ_1 . For this we place the surface of integration around

⁹Dissertation, Breslau 1909.

¹⁰Ann. d. Phys. **23**, 626 (1908).

¹¹Proc. Tokyo Mathem. Physical Soc. (2) 4, 966 (1908).

¹²Kirchhoff, Lectures on Mathematical Physics, Vol. II, Optics, 1891 (in German).

or finally,

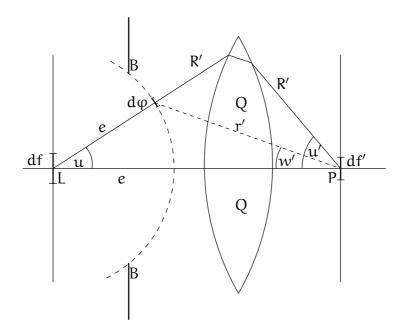
$$\rho = -\frac{x\xi + y\eta}{e} \,. \tag{17}$$

This simplification of the value ρ for z = -e, i.e., for the observation points *that lie in the object plane itself*, acquires a *physical* meaning with the introduction of imaging systems.

§13. Diffraction phenomena occurring in pairs of conjugate planes of optical systems

In Fig. 20, let the surface element df lying at L glow and its image df', projected by system Q, lie at P. Let diaphragm BB act as the entrance pupil that cuts an effective piece of the surface out of a





which, since e' is large compared to λ' , reduces to the expression identical to Eq. 26,

$$\mathfrak{E}' = rac{\mathrm{const}}{\mathbf{e}'} \cos \mathfrak{u}' \cdot \cos 2\pi \left(rac{\mathrm{t}}{\mathrm{T}} + rac{\mathbf{e}'}{\lambda'}
ight) \; .$$

With this it has been shown that \mathfrak{E}' is a solution of the wave equation for the case treated here and therefore can be inserted in place of φ in Eq. 13 of the Kirchhoff principle.

If one introduces once again s' via Eq. 24a,

$$J_{P_1} = \overline{\mathfrak{E}'^2} = \overline{s'^2} \cdot \overline{df'} ,$$

after easy calculation, ^{xli} if one replaces r with e' in the amplitude and $\frac{1+\cos u}{2}$ with 1, one obtains

$$\begin{split} s' &= \frac{k'}{\lambda'} \int_{II} \frac{d\phi' \cos u'}{e'^2} \sin 2\pi \left(\frac{t}{T} + \frac{x'\xi' + y'\eta'}{e'\lambda'} \right) \\ &= \frac{k'}{\lambda'} \int_{II} \frac{d\xi' \, d\eta'}{e'^2} \sin 2\pi \left(\frac{t}{T} + \frac{x'\xi' + y'\eta'}{e'\lambda'} \right) \,, \end{split}$$

which is exactly the above derived expression (Eq. 24).

It should be pointed out once more that one obtains the "effective piece of boundary surface I" as one draws from the luminous point or surface element all possible rays toward the boundary points on the entrance pupil. The entirety of the intersections of these rays with the spherical surface I form the boundary of the "effective piece." Integration in the expression of s is extended over the projection of this "effective piece" onto the $\xi\eta$ -plane.

§17. Calculation of diffraction on an aperture of specific form for points in the plane conjugate to the object plane in the presence of a luminous surface element

We choose the form of the diffracting aperture in such a way that *the projection of the effective piece of the boundary surface onto the* $\xi\eta$ *-plane is*

Imaging of illuminated objects

§18. Presence of several luminous points

In the presence of *one* luminous surface element, the diffraction pattern is symmetrical with respect to the location of that element. This applies to an arbitrarily located surface element, as long as one limits oneself to points close to the axis of the system. *The diffraction pattern always remains stationary and moves with the luminous surface element*.

With the simultaneous presence of several luminous elements, the observed diffraction pattern depends on whether the individual elements emit independent *incoherent* waves from each other, or whether the waves emitted from individual elements are *coherent*, *i.e.*, *capable of interference*.

The following laws hold, assuming that we are dealing with several luminous "points": If different wave trains are incoherent, one obtains the resulting intensity at each location by simply summing the squares of the amplitudes, i.e., the intensities, that are generated by individual luminous points.

If n luminous "points" contribute to the light disturbance at the observation point, and if the disturbance generated by their wave trains are represented by the value of the electric field (of the light vector),

$$A(\mathbf{x}) = \frac{2}{\pi} \int_{0}^{\frac{2\pi\alpha\alpha'}{\lambda}} dw \cos\left(\frac{\mathbf{x}w}{\alpha}\right)$$
$$= \frac{2\alpha}{\pi \mathbf{x}} \cdot \sin\left(\frac{2\pi\alpha'\mathbf{x}}{\lambda}\right)$$
$$= \frac{4\alpha\alpha'}{\lambda} \cdot \frac{\sin\left(\frac{2\pi\alpha'\mathbf{x}}{\lambda}\right)}{\frac{2\pi\alpha'\mathbf{x}}{\lambda}}.$$
(56)

A(x) has in this case the already discussed form $\frac{\sin w}{w}$. If $\frac{2\pi a \alpha'}{\lambda}$ is very large compared to π , then, as can be seen from the consideration of the form of $\frac{dA(x)}{dx}$, the fluctuations of the amplitude inside the slit are very small, and the value of the amplitude is therefore almost constant; only *at the edges* of the slit do fluctuations take place; namely (if we consider only positive values of x, since the phenomenon is symmetrical with respect to the J-axis), since $\frac{2\pi a \alpha'}{\lambda}$ was already assumed to be large, u is a fortiori large and therefore:

$$\frac{\mathrm{d}A(\mathbf{x})}{\mathrm{d}\mathbf{x}} = -\mathrm{const}\cdot\frac{\sin\nu}{\nu} \ .$$

Therefore, as v gets closer and closer to the value v = 0 (as x increases), i.e., x = a (edge of the slit), the fluctuations of $\frac{\sin v}{v}$ begin to become more and more noticeable. We therefore obtain the image of the amplitude indicated in Fig. 42:^{xlix} the larger $\frac{a\alpha'}{\lambda}$ becomes, the more the variations at the edges converge, so that in the limit, for infinitely large $\frac{a\alpha'}{\lambda}$, we obtain the amplitude graph already shown in Fig. 37 above.

§23. Finite slit whose two halves possess a constant difference in phase

Let the slit have width 2a and height 2b; let the phase in the half slit of height 2b and width a (x = -a to x = 0) be equal to $2\pi \frac{t}{T}$, while

95

to the grating constant γ . The larger γ becomes, the more diffraction maxima can contribute to image formation, and the greater the similarity. The maximum numerical aperture of a system is reached when $U = 90^{\circ}$ and is then

$$A = n$$
.

Therefore, in this case of maximum possible performance,

$$h = n \frac{\gamma}{\lambda_0} \,. \tag{79}$$

If we denote with h_1 the last diffraction spectrum of intensity or brightness to be considered in the overall image of the function $f(\xi', \eta')$, the system with A = n will image all gratings with absolute similarity, if

$$\gamma \geqslant \frac{h_l \cdot \lambda_0}{n}$$

§27. Dissimilar imaging of the object

We shall base this investigation on a system with maximum aperture A = n, which still images a grating with constant γ with absolute similarity, meaning the satisfaction of the inequality

$$\gamma \geqslant h_l \lambda_0 / n$$
 ,

where h_l is the last diffraction spectrum of intensity still to be considered in the overall image of the function $f(\xi', \eta')$. A grating with a smaller grating constant ($\gamma' < \gamma$) is therefore no longer imaged by the system similarly. If λ_0 has the smallest possible value (photographic waves) and n has the highest possible value (homogeneous immersion), then the grating $\gamma = h_l \lambda_0 / n$ is imaged in an absolutely similar way (a fortiori all gratings with *larger* grating constants), whereas it is physically impossible to image gratings with smaller grating constants ($\gamma' < \gamma$) similarly.

As an example, let us suppose that $\lambda_0 = 350 \text{ nm}$, n = 1.65, and $h_1 = 10$, assuming that maxima with an intensity less than 1 % of the

Imaging of a grating with artificial clipping of diffraction orders¹

§28. General intensity equation

Finally, as a typical example, we want to treat the imaging of a grating. Let the grating extend along the X-axis from X = -A to X = +A, and along the Y-axis from Y = -B to Y = +B, so that it lies symmetrically with respect to the X- and Y-axis and let it consist of N slits of width 2a, which are separated by "bars" of width 2Δ . Therefore, $\gamma = 2(a + \Delta)$ is the grating constant. Let N be a large number. Let α' and β' be the angular height and width of the diffracting aperture (boundary), which lies as a whole or in its parts symmetrically to the X- and Y-axis.

¹The results given in this chapter are taken, at our urging, from the doctoral dissertation of M. Wolfke (Breslau 1910), which will soon appear in *Annalen der Physik*.

However, if $\frac{2a}{\gamma} > 0.6$, then the minima of A are positive and yield, after squaring, minima in intensity.

The decrease in intensity from the maximum to the minimum is in the form of a cosine, for it follows the law $I = (1 + C \cos u)^2$, where $u = \frac{2\pi x}{\gamma}$; maxima and minima have equal width (see Figs. 57a and b). We therefore obtain the following result:

If, in addition to the central order, the first two side maxima also contribute to the secondary image, the image shows a structure. The number of grating lines is reproduced correctly in the image, but the intensity drop from the maximum to the minimum is gradual, and the maxima and minima appear equally wide. In addition, under certain circumstances, secondary maxima still occur in the middle of the minima.

§31. Case III: Only the ith maxima on both sides contribute to imaging; the central image is blocked

The expression for intensity now becomes

$$I = \text{const} \cdot 4 \left[\int_{\frac{2\pi a(Ni+1)}{N\gamma}}^{\frac{2\pi a(Ni+1)}{N\gamma}} dw \frac{\sin w \sin \frac{N\gamma w}{2a}}{w \sin \frac{\gamma w}{2a}} \cos \frac{x}{a} w \right]^{2} \right\} .$$
(96)
= const \cdot J_{i}^{2}

If we introduce a new variable,

$$w_1 = \pi \mathrm{i} - rac{\gamma w}{2 \mathfrak{a}}$$
 ,

for the transformation of J_i , the integration limits will become symmetrical with respect to the origin; we can then again omit, as in case II above, the integral over the odd function and finally obtain, after introducing

$$w_2 = \frac{2a}{\gamma}w_1$$

A brief introduction to geometrical optics

Fermat's Principle: Geometrical optics deals with the (artificial) concept of light rays. A light ray from point P to point P' is a P- and P'-containing path s that is always perpendicular to the successive wavefronts of the light as it propagates from P and P'. The optical length, which is the cumulative phase, is then equal to $\int_{P}^{P'} n \, ds$, where n is the index of refraction along the path. Being perpendicular to the two neighboring wavefronts, ds is the shortest distance between them. Therefore, if l is any other path connecting points P and P', it must be that

$$\int_{P}^{P'} n \, ds \leqslant \int_{P}^{P'} n \, dl$$

This is the same as stating that the optical length $\int_{P}^{P'} n \, ds$ from points P to P' is a stationary one. This is called Fermat's principle. To find the actual ray path, we start by considering an arbitrary path from P to P', vary it while holding the two ends fixed, and set the variation $\delta(\int_{P}^{P'} n \, dl)$ to zero.



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